# Moving and jumping spot in a two dimensional reaction-diffusion model 

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#### Abstract

We consider a single spot solution for the Schnakenburg Model in a two-dimensional unit disk in the singularly perturbed limit of a small diffusivity ratio. For large values of the reaction-time constant, this spot can undergo two different types of instabilities, both due to a Hopf bifurcation. The first type induces oscillatory instability in the height of the spot. The second type induces a periodic motion of the spot center. We use formal asymptotics to investigate when these instabilities are triggered, and which one dominates. In the parameter regime where spot motion occurs, we construct a periodic solution consisting of a rotating spot, and compute its radius of rotation and angular velocity. Detailed numerical simulations are performed to validate the asymptotic theory, including rotating spots. More complex, non-circular spot trajectories are also explored numerically.


## 1. INTRODUCTION

The Schnakenburg model, introduced in [1], is a particular case of the activator-substrate system. It was originally formulated as a simplified model of a trimolecular autocatalytic reaction with diffusion. It is also a limiting case of both the Gray-Scott model as well as the Klausmeyer model of vegetation pattern formation on a flat ground when the water evaporation is limited [2]. Some applications of Schankenburg model to biology include pattern formations in embryogenesis and skin patterns [3, 4]. However arguably its biggest value is that it serves as a simple prototype model for studying pattern formation in reaction-diffusion systems: it is among the simplest class of models which generate stable inhomogeneous patterns. As such, the Schnakeberg and related models such as Gray-Scott, the Brusselator, and the Gierer-Meinhardt model, have been extensively studied (especially in one, but also and two and higher dimensions), and phenomena such as spike formation, stability, self-replication, oscillations and motion has been analysed in detail. A very incomplete list of references includes [5-18].

In [7-9] the authors have found that a single spike in a one-dimensional Gray-Scott model can undergo destabilizing oscillations in either its height or position. The height oscillations happen on a much faster timescale when compared to the position oscillations. Which instability is triggered first depends on the value of the feed-rate $A$ representing the amount of the substrate chemical that is being pumped into the system. Typically, height oscillations were triggered at lower feed rates than the position oscillations. Periodic spike motion in one dimensional GS model was further investigated in [17].

In one-dimensional domain, there has been much work over the past decade in analyzing the stability, dynamics, and self-replication of spike patterns for the Schakenberg model and related models with similar structure. The stability problem for equilibrium spike patterns in infinite domain has been studied in [14] and [11] following earlier work on Gierer-Meinhardt model [19]. In [20], the authors studied Hopf bifurcations and oscillatory instabilities of spike solutions of Gierer-Meinhardt model for various ranges of the reaction-time constant. For a recent summary on pattern formation in GM model, see [21] and references therein. A detailed study of self-replication, overcrowding instability, and spike height and position oscillations for the Gray-Scott model is conducted in [6-9, 17, 22]. Selfreplication in slowly growing domains was also studied in [23] and [24].

In two dimensions, Muratov and Osipov [10] were among the first to study the Gray-Scott model, including selfreplication thresholds. Wei and Winter [12] reviewed analytical methods for a rigorous study of the existence and stability of stationary, multiple spots for reaction-diffusion systems and considered two classes of reaction-diffusion systems: activator-inhibitor systems (such as the Gierer-Meinhardt system) and activator-substrate systems (such as the Gray-Scott system or the Schnakenberg model). In [13], spot replication for the Schnakenberg model was studied. In [16], the authors studied multi-spot patterns including competition, spot motion, and self-replication, for the related Gray-Scott model.

The purpose of this work is two-fold. First, we extend the results in [7-9] on thresholds for oscillatory instabilities for both height and spike position (and particularly periodically moving spikes) from one dimension to a two-dimensional setting. Second, we investigate the spot trajectory (i.e. the path traced by the spot center) and the kind of complex two-dimensional motion that can result.

We will use the following scaling of the Schankenberg model,

$$
\begin{equation*}
v_{t}=\varepsilon^{2} \Delta v-v+v^{2} u, \quad \tau u_{t}=\Delta u+A-\frac{v^{2} u}{\varepsilon^{2}} \quad \text { inside } \Omega \subset \mathbb{R}^{2} ; \quad \partial_{n} v=\partial_{n} u=0 \quad \text { on } \partial \Omega \tag{1.1}
\end{equation*}
$$

Throughout this paper, we assume $\Omega$ to be a unit disk,

$$
\begin{equation*}
\Omega \text { is a unit disk; } \Omega=\left\{x \in \mathbb{R}^{2}:|x|<1\right\} \tag{1.2}
\end{equation*}
$$



FIG. 1. Two types of dynamics of a single spike solution of the Schnakenberg model. Top row: height oscillations on $O(1)$ timescale. Three snapshots of $v(x, t)$ are shown at times as indicated. Parameters are $\varepsilon=0.03, A=1, \tau=0.07$. Top right shows the height of the spike as a function of time. The spike remains at the center of the disk. Bottom row: periodic motion of the spot on a slow timescale. Three snapshots of $u(x, t)$ are shown. Bottom right shows the trajectory of the spot center. Parameters are $\varepsilon=0.02, A=8, \tau=0.15 / \varepsilon^{2}$. Dashed line shows the asymptotic prediction for the spot trajectory (Proposition 6.1)
although some of our results can be extended to more general domains.
Here, $\varepsilon \ll 1, A>0$ and $\tau>0$, represent diffusivity, the feed-rate and the reaction-time constant respectively. The equations model the following process: the fast-diffusing substrate $u$ is consumed by a slowly diffusing activator $v$, which decays in time. The substrate is being pumped into the system at a constant rate, represented by parameter $A$. The reaction kinetics for $u$ and $v$ occur at different scales (depending on the choice of $\tau$ ). Of particular interest to us will be the regime where $\tau$ is very large, so that $u$ reacts much slower than $v$. As we will show, the oscillatory instabilities (both for spike height and positions) are triggered when $\tau$ is very large.

In this paper we consider the effect of increasing the parameter $\tau$ on a single spot at the center of the unit disk. The associated linearized eigenvalue problem has eigenfunctions of the form $\phi(r) e^{i m \theta}$ in the polar coordinates. Due to underlying translational invariance, the eignvalues corresponding to mode $m= \pm 1$ are asymptotically small as $\varepsilon \rightarrow 0$ and their instability induces a slow (possibly periodic) motion of the spot. We refer to these eigenvalues as small eigenvalues. All other eigenvalues are referred to as large eigenvalues. The mode $m=0$ corresponds to purely radial perturbations and its instability can induce spike oscillation or collapse, whereas the instability with respect to mode $m=2$ eigenvalues triggers self-replication [10, 13, 16, 25]. Here, we only concentrate on modes $m=0,1$ since $\tau$ does not appear to trigger instability of the higher modes.

Two different types of instabilities can be triggered when $\tau$ is sufficiently increased as illustrated in Figure 1: either large or small eigenvalues can undergo a Hopf bifurcation. The former instability triggers height oscillations, whereas the latter triggers slow translational instabilities in spike position, inducing (typically periodic) spike motion. Which one is triggered first depends on values of $\varepsilon$ and $A$.

Our main task is to classify precisely for which parameters $A$ and $\varepsilon$ does the spike motion (as opposed to height oscillations) occur when $\tau$ is increased sufficiently? In §5 (Proposition 5.1) we give a concise characterisation in terms of the following threshold. Let

$$
\begin{equation*}
A_{c} \sim \frac{6.283}{\sqrt{\log \left\{\left(\log \frac{1}{\varepsilon}\right) 1.010-0.1433\right\}}} \tag{1.3}
\end{equation*}
$$


(a)

| $\varepsilon$ | $A_{c}$ (numerics) | $A_{c}$ (theory) |
| :--- | :---: | :---: |
| 0.02 | 6.264 | 5.4340 |
| 0.01 | 5.835 | 5.1204 |
| 0.005 | 5.529 | 4.8910 |
| 0.0025 | 5.293 | 4.7140 |
| 0.00125 | 5.029 | 4.5724 |

(b)

FIG. 2. (a) Hopf bifurcations values for height and position oscillations ( $\tau_{h, \text { large }}$ and $\tau_{h, \text { small }}$, respectively), as a function of $A$, and with $\varepsilon=0.02$. Theoretical predictions given by Propositions 3.1 and 4.1 are also shown. The intersection of $\tau_{h, l a r g e}$ and $\tau_{h, \text { small }}$ is denoted by $A_{c}$. (b) Comparison between the asymptotic value of $A_{c}$ given by (1.3) and the numerically computed value for several small $\varepsilon$.

In the limit $\varepsilon \rightarrow 0$, for values of $A$ bigger than $A_{c}$, spike motion is observed as $\tau$ is sufficiently increased, whereas when $A<A_{c}$, increasing $\tau$ triggers height oscillations. This threshold follows from computing the thresholds $\tau_{h, \text { large }}$ and $\tau_{h, \text { small }}$, corresponding to the Hopf bifurcation points for small and large eigenvalues, respectively. This is done in $\S 3$, and $\S 4$ to obtain

$$
\begin{equation*}
\tau_{h, \text { large }} \sim \frac{19.929}{A^{2} \varepsilon^{2}} \exp \left(\frac{-39.474}{A^{2}}\right) ; \quad \tau_{h, \text { small }} \sim \frac{1}{\varepsilon^{2} A^{2}} \frac{19.737}{\log \frac{1}{\varepsilon}-0.1419} \tag{1.4}
\end{equation*}
$$

in the critical regime $O\left(\frac{1}{\log 1 / \varepsilon}\right) \ll A^{2} \ll O(1)$. In fact the threshold (1.3) is obtained from (1.4) by simply setting $\tau_{h, \text { large }}=\tau_{h, \text { small }}$.

The threshold (1.3) has a striking log log scaling. Despite such a slow convergence rate, remarkably it agrees relatively well with numerical experiments even when $\varepsilon=0.01$ (see Figure 2).

In $\S 6$ we analyse what happens beyond the Hopf bifurcation for small eignevalues, in the regime $\tau_{h, s m a l l}<\tau<$ $\tau_{h, \text { large }}$. In this regime, the spot starts to move and there exists time-dependent solutions in a form of a rotating spot. We compute the radius and speed of the rotation in Proposition 6.1. By computing the radius of rotation $r_{0}$ as a function of $\tau$, we find that $r_{0} \rightarrow 0$ as $\tau \rightarrow \tau_{h, s m a l l}$ from above. In other words, the rotating spot solution bifurcates from a stationary spot as a result of a Hopf bifurcation. This is illustrated in Figure 4. We conclude with numerical experiments demonstrating even more complex spike motion (see Figure 5).

## 2. EQUILIBRIUM SOLUTION

We start by reviewing the construction of the equilibrium solution to (1.1) using the method of matched asymptotic expansions as was previously done in [13]. At the equilibrium, the steady state satisfies

$$
\begin{equation*}
0=\varepsilon^{2} \Delta v-v+u v^{2}, \quad 0=\Delta u+A-\frac{u v^{2}}{\varepsilon^{2}} \tag{2.5}
\end{equation*}
$$

with Neumann boundary conditions on $\Omega$. We assume that $\Omega$ is a unit disk with the spike located at the center. Near the core of the spike, we rescale:

$$
\begin{equation*}
v(x)=V(y), \quad u(x)=U(y) \quad, y=\varepsilon^{-1} x \tag{2.6}
\end{equation*}
$$

Then (2.5) becomes

$$
\Delta_{y} V-V+U V^{2}=0, \quad \Delta_{y} U+A \varepsilon^{2}-U V^{2}=0, \quad y \in \mathbb{R}^{2}
$$

where $\Delta_{y}$ denotes the Laplacian in $y$. We expand

$$
U=U_{0}+\varepsilon^{2} U_{1}+\cdots, \quad V=V_{0}+\varepsilon^{2} V_{1}+\cdots
$$

To leading order, we look for a radially symmetric solution given by $V_{0}=V_{0}(\rho)$ and $U_{0}=U_{0}(\rho)$, with $\rho=|y|$. It satisfies the following coupled nonlinear radially symmetric "core problem",

$$
\begin{gather*}
\Delta_{\rho} V_{0}-V_{0}+U_{0} V_{0}^{2}=0, \quad \Delta_{\rho} U_{0}-U_{0} V_{0}^{2}=0, \quad 0<\rho<\infty  \tag{2.7a}\\
V_{0} \rightarrow 0, \quad U_{0} \sim S \log \rho+\chi(S) \quad \text { as } \rho \rightarrow \infty \tag{2.7~b}
\end{gather*}
$$

The core problem (2.7), was first identified in one dimension in [5]. It is closely related to the phenomenon of self-replicating spots $[10,13,16]$.

To determine the source strength $S$, we integrate the second equation in (2.5) to obtain

$$
A \pi=\int_{\Omega} \frac{u v^{2}}{\varepsilon^{2}} d x \sim \int_{\mathbf{R}^{2}} U_{0} V_{0}^{2} d y
$$

On the other hand integrating the second equation in (2.7a) and using the divergence theorem, we obtain

$$
2 \pi S=\int_{\mathbf{R}^{2}} U_{0} V_{0}^{2} d y
$$

so that

$$
\begin{equation*}
S=\frac{A}{2} \tag{2.8}
\end{equation*}
$$

In general, the solution to (2.7a) as well as the function $\chi(S)$ in (2.7b) must be computed numerically. This was done for example in $[10,13,16]$. Figure 2(b) shows the function $\chi(S)$. However for small $S$, equations (2.7a) become weakly coupled since $U_{0}$ becomes nearly constant and we may estimate the solution to (2.7a) as follows. Assume that $S \ll 1$ and $U_{0}(y) \sim U_{0}$ is constant to leading order in $S$. Then $V_{0}, U_{0}$ satisfy at leading order,

$$
\begin{equation*}
V_{0}(y)=w(y) \sigma ; \quad U_{0}(y)=1 / \sigma \tag{2.9}
\end{equation*}
$$

where $w$ is the unique positive ground-state solution to

$$
\begin{equation*}
\Delta w-w+w^{2}=0 ; \quad w \rightarrow 0 \text { as }|y| \rightarrow \infty \tag{2.10}
\end{equation*}
$$

and

$$
2 \pi S \sim \sigma \int w^{2} d y ; \quad \chi(S) \sim \frac{1}{\sigma}
$$

This yields an asymptotic expression

$$
\begin{equation*}
A \sim 2 \sigma \int_{0}^{\infty} w^{2}(\rho) \rho d \rho, \quad S \sim \sigma \int_{0}^{\infty} w^{2}(\rho) \rho d \rho, \quad \chi(S) \sim \frac{1}{\sigma} \tag{2.11}
\end{equation*}
$$

The resulting integral is evaluated numerically (see Appendix B). We summarize this construction as follows.


FIG. 3. (a) Core problem $V_{0}, U_{0}$ ) and its asymptotics (Proposition 2.1). Good comparison between numerics and asymptotics (2.9) is observed even for relatively "large" $A=2$. (b) $A$ versus $\chi$. Asymptotics denote the regime $A \ll O(1)$ given by $\chi \approx 9.868 / A$.

Proposition 2.1 In the limit $0<\varepsilon \ll 1$, the leading order steady-state solution to (2.5) near the origin satisfies

$$
\begin{equation*}
v(x) \sim V_{0}(y), \quad u(x) \sim U_{0}(y) \quad, y=\varepsilon^{-1} x \tag{2.12}
\end{equation*}
$$

where $V_{0}, U_{0}$ satisfy core problem (2.7) and where the constant $S$ is given by (2.8). In the regime $0<A \ll O(1)$, we have the asymptotics

$$
\begin{align*}
V_{0}(x) & \sim \sigma w(y), \quad y=\varepsilon x  \tag{2.13a}\\
U_{0}(x) & \sim \frac{1}{\sigma} \tag{2.13b}
\end{align*}
$$

where $w(y)$ is the unique ground state given by (2.10) and where

$$
\begin{equation*}
\sigma=\frac{1}{2 \int_{0}^{\infty} w^{2} \rho d \rho} A \approx \frac{A}{9.868} \ll 1 \tag{2.13c}
\end{equation*}
$$

Figure 2(b) shows the graph of $S$ versus $\chi(S)$, as well as its asymptotic approximation. While the asymptotics are formally valid in the regime $A \ll 1$, they agree well with full numerics even when $A$ is relatively large. For example when $A=2$, full numerics yield $\chi \approx 5.103$ whereas formula ( 2.11 ) yields $\sigma \approx 0.2026, \quad \chi \approx 4.934$, for a relative error of only $3 \%$. Note also that this error is independent of $\varepsilon$ to leading order. Figure 2(a) gives a comparison between $V_{0}(x)$ and $\sigma w(y)$ with $A=2$. Excellent agreement is observed. This is in part because the effective small parameter is $\sigma \approx 0.2 \ll 1$ when $A=2$.

## 3. STABILITY: LARGE (MODE ZERO) EIGNEVALUE

Having constructed the steady state, we now consider its stability. Linearizing around the steady state in (1.1) we write.

$$
u(x, t)=v(r)+e^{\lambda t} \phi(x), \quad u(x, t)=u(r)+e^{\lambda t} \psi(x)
$$

The linearized system of (1.1) then becomes:

$$
\left\{\begin{array}{cc}
\lambda \phi=\varepsilon^{2} \Delta \phi-\phi+v^{2} \psi+2 u v \phi & x \in \Omega  \tag{3.14}\\
\tau \lambda \psi=\Delta \psi-\frac{1}{\varepsilon^{2}}\left(v^{2} \psi+2 u v \phi\right) & \\
\partial_{n} \phi=0=\partial_{n} \psi & x \in \partial \Omega
\end{array}\right.
$$

In the inner region, we expand

$$
\begin{equation*}
\phi=e^{i m \theta}\left(\Phi_{0}(\rho)+\varepsilon^{2} \Phi_{1}(\rho)+\cdots\right) \quad \psi=e^{i m \theta}\left(\Psi_{0}(\rho)+\varepsilon^{2} \Psi_{1}(\rho)+\cdots\right) \quad \rho=|y| \quad y=\varepsilon^{-1} x \tag{3.15}
\end{equation*}
$$

where $y=\rho(\cos \theta, \sin \theta)$. Substituting (3.15) into (3.14), then to leading order we obtain the following radially symmetric eigenvalue problem:

$$
\begin{align*}
\lambda \Phi_{0} & =\Delta_{m} \Phi_{0}-\Phi_{0}+V_{0}^{2} \Psi_{0}+2 U_{0} V_{0} \Phi_{0} \\
0 & =\Delta_{m} \Psi_{0}-\left(V_{0}^{2} \Psi_{0}+2 U_{0} V_{0} \Phi_{0}\right) \tag{3.16}
\end{align*}
$$

Here $\Delta_{m} \Phi_{0} \equiv \partial_{\rho \rho} \Phi_{0}+\rho^{-1} \partial_{\rho} \Phi_{0}-m^{2} \rho^{-2} \Phi_{0}$. and $U_{0}, V_{0}$ are solutions to (2.7).
Because of the decay term in the equation for $\Phi_{0}$, we assume that $\Phi_{0}$ decays exponentially for large $|y|$. On the other hand, the appropriate far-field boundary condition for $\Psi_{0}$ depends on whether $m=0$ or $m \geqq 1$.

We begin by considering the mode $m=0$. In this case the far-field conditions for $\Psi_{0}$ exhibits logarithmic growth, $\Psi_{0} \sim C \log |y|+B,|y| \gg 1$. We can scale the eigenfunction to set $C=1$, so that $\Psi_{0}$ then satisfies

$$
\begin{equation*}
\Psi_{0} \sim \ln |y|+B, \quad|y| \gg 1 \tag{3.17}
\end{equation*}
$$

By integrating over the equation for $\Psi_{0}$ and using the Divergence theorem, this scaling is equivalent to

$$
\begin{equation*}
\int_{\mathbb{R}^{2}}\left(V_{0}^{2} \Psi_{0}+2 U_{0} V_{0} \Phi_{0}\right)=2 \pi \tag{3.18}
\end{equation*}
$$

The constant $B$ is determined by matching to the outer region. Since $v$ is assumed to decay away from the spike, using (3.18) we have

$$
\int_{\Omega} \frac{1}{\varepsilon^{2}}\left(v^{2} \psi+2 u v \phi\right) d x \sim 2 \pi
$$

and the outer problem for $\psi$ is

$$
\begin{equation*}
\lambda \tau \psi=\Delta \psi-2 \pi \delta(x) \text { inside } \Omega ; \quad \partial_{n} \psi=0 \text { on } \partial \Omega \tag{3.19}
\end{equation*}
$$

whose solution is given by

$$
\begin{equation*}
\psi(x) \sim \frac{K_{0}^{\prime}(\sqrt{\tau \lambda})}{I_{0}^{\prime}(\sqrt{\tau \lambda})} I_{0}(\sqrt{\tau \lambda} r)-K_{0}(\sqrt{\tau \lambda} r), \quad r=|x| . \tag{3.20}
\end{equation*}
$$

Expanding for small $r$ we have

$$
\begin{align*}
\psi(x) & \sim \log (r)+\frac{K_{0}^{\prime}(\sqrt{\tau \lambda})}{I_{0}^{\prime}(\sqrt{\tau \lambda})}-\log (2)+\gamma+\log (\sqrt{\lambda \tau}) \quad \text { as } r \rightarrow 0 \\
& \sim \log (|y|)+\frac{K_{0}^{\prime}(\sqrt{\tau \lambda})}{I_{0}^{\prime}(\sqrt{\tau \lambda})}+\log \left(\frac{e^{\gamma}}{2} \sqrt{\varepsilon^{2} \lambda \tau}\right) \tag{3.21}
\end{align*}
$$

Note that the above expansion assumes that $\sqrt{\tau \lambda} \varepsilon \ll 1$. This will be shown to be self-consistent later on. Then matching (3.17) and (3.21), yields

$$
\begin{equation*}
B=\frac{K_{0}^{\prime}(\sqrt{\tau \lambda})}{I_{0}^{\prime}(\sqrt{\tau \lambda})}+\log \left(\frac{e^{\gamma}}{2} \sqrt{\varepsilon^{2} \lambda \tau}\right) \tag{3.22}
\end{equation*}
$$

Together with (3.16), this provides a closed-system which determines the eigenvalue $\lambda$. We summarize this construction:

Proposition 3.1 In the limit $\varepsilon \rightarrow 0$, the mode-zero eigenvalue $\lambda$ of the lineraized problem (3.14) is asymptotic to the eigenvalue problem (3.16) with $m=0$ subject to the outer condition (3.17) where $B$ is given by (3.22), as long as $\sqrt{\tau \lambda} \varepsilon \ll 1$.

We now concentrate on the weakly-coupled regime to $A \ll 1$ given by (2.13c). Substitute the steady-state expansion (2.13) into (3.16) to obtain to leading order

$$
\begin{align*}
\lambda \Phi_{0} & =\Delta_{0} \Phi_{0}-\Phi_{0}+2 w \Phi_{0}+w^{2} \sigma^{2} \Psi_{0}  \tag{3.23}\\
0 & =\Delta_{0} \Psi_{0}-\left(w^{2} \sigma^{2} \Psi_{0}+2 w \Phi_{0}\right)
\end{align*}
$$

Rescale $\Phi_{0}=\sigma^{2} \hat{\Phi}_{0}$ and drop the hat to obtain

$$
\begin{gather*}
\lambda \Phi_{0}=\Delta_{0} \Phi_{0}-\Phi_{0}+2 w \Phi_{0}+w^{2} \Psi_{0}  \tag{3.24}\\
0=\Delta_{0} \Psi_{0}-\sigma^{2}\left(w^{2} \Psi_{0}+2 w \Phi_{0}\right)
\end{gather*}
$$

whereas (3.18) becomes

$$
\begin{equation*}
\int_{\mathbb{R}^{2}}\left(w^{2} \Psi_{0}+2 w \Phi_{0}\right) d y=2 \pi \sigma^{-2} \tag{3.25}
\end{equation*}
$$

In addition, we will assume à-priori that $|\tau \lambda| \gg 1$ (this self-consistency of this assumption will be verified at the end). Under this assumption, using the large-argument expansion of the Bessel functions, the term $\frac{K_{0}^{\prime}(\sqrt{\tau \lambda})}{I_{0}^{\prime}(\sqrt{\tau \lambda})}$ is exponentially small so that

$$
B \sim \log \left(\frac{e^{\gamma}}{2} \sqrt{\varepsilon^{2} \lambda \tau}\right)
$$

Furthermore suppose $B \gg 1$. Then we may estimate $\Psi_{0}$ by a constant,

$$
\Psi_{0} \sim \log \left(\frac{e^{\gamma}}{2} \sqrt{\varepsilon^{2} \lambda \tau}\right)
$$

We further rescale $\Phi_{0}(y)=-\Psi_{0} \Phi(y)$ which leads to the reduced problem

$$
\begin{align*}
\left(L_{0}-\lambda\right) \Phi & =w^{2}  \tag{3.26a}\\
-2 \int w \Phi+\int w^{2} d x & =\frac{2 \pi \sigma^{-2}}{\log \left(\frac{e^{\gamma}}{2} \sqrt{\varepsilon^{2} \lambda \tau}\right)} \tag{3.26b}
\end{align*}
$$

where the operator $L_{0}$ is defined by

$$
\begin{equation*}
L_{0} \Phi:=\Delta_{0} \Phi-\Phi+2 w \Phi \tag{3.27}
\end{equation*}
$$

One of the key properties of the operator $L_{0}$ is that

$$
L_{0} w=w^{2}
$$

as can be readily verified using (2.10). This suggests that we seek a Hopf bifurcation point of (3.26) assuming $\lambda$ is small. We therefore expand in $\lambda$

$$
\begin{equation*}
\Phi=w+\lambda \Phi_{1}, \quad \lambda \ll 1 \tag{3.28}
\end{equation*}
$$

to obtain

$$
\Phi_{1}=L_{0}^{-1}(w)
$$

Define

$$
\begin{equation*}
\tau_{0}:=\left(\frac{e^{\gamma}}{2}\right)^{-2} \varepsilon^{2} \tau \tag{3.29}
\end{equation*}
$$

and assume that $\lambda$ is purely imaginary,

$$
\lambda=i \lambda_{I} ; \quad \lambda_{I} \ll 1
$$

Then (3.26) becomes

$$
\begin{equation*}
-2 \lambda_{I} i \int w L_{0}^{-1}(w) d y-\int w^{2} d y=\frac{2 \pi \sigma^{-2}}{\log \left(\sqrt{i \lambda_{I} \tau_{0}}\right)} \tag{3.30}
\end{equation*}
$$

Using the identity

$$
L_{0}^{-1} w=w+\frac{1}{2} y \cdot \nabla w=w+\frac{1}{2} \rho w^{\prime}(\rho)
$$

and integrating by parts, we obtain

$$
\begin{equation*}
2 \int w L_{0}^{-1}(w) d y=\int w^{2} d y \tag{3.31}
\end{equation*}
$$

so that (3.30) becomes

$$
\begin{equation*}
\log \left(\sqrt{i \lambda_{I} \tau_{0}}\right)=\frac{-2 \pi \sigma^{-2}}{\int w^{2} d y} \frac{1}{\lambda_{I} i+1}=\frac{2 \int w^{2} d y}{A^{2} \pi}\left(\frac{i \lambda_{I}-1}{\lambda_{I}^{2}+1}\right) \tag{3.32}
\end{equation*}
$$

Equating real and imaginary parts we obtain

$$
\left\{\begin{array}{c}
\frac{1}{2} \log \left(\lambda_{I} \tau_{0}\right)=\frac{2 \int w^{2} d y}{A^{2} \pi} \frac{-1}{\lambda_{I}^{2}+1}  \tag{3.33}\\
\frac{\pi}{4}=\frac{2 \int w^{2} d y}{A^{2} \pi} \frac{\lambda_{I}}{\lambda_{I}^{2}+1}
\end{array}\right.
$$

These equations yields, to leading order in $A \ll 1$,

$$
\left\{\begin{array}{c}
\lambda_{I} \sim \frac{\pi^{2} A^{2}}{8 \int w^{2} d y}  \tag{3.34}\\
\tau_{0}=\exp \left(\frac{-4 \int w^{2} d y}{A^{2} \pi}\right) \frac{8 \int w^{2} d y}{\pi^{2} A^{2}}
\end{array}, A \ll 1\right.
$$

Using (3.29) we finally obtain the critical value of $\tau=\tau_{h}$ at the Hopf bifurcation point for large eigenvalue:

$$
\tau_{h}=\frac{1}{A^{2} \varepsilon^{2}} \exp \left(\frac{-4 \int w^{2} d y}{A^{2} \pi}\right) \frac{2 e^{2 \gamma} \int w^{2} d y}{\pi^{2}}
$$

We made three assumptions in this derivation: (i) $\lambda \tau \gg 1$; (ii) $\varepsilon^{2} \tau \lambda \ll 1$ and (iii) $\lambda \ll 1$. Assumptions (ii) and (iii) are satisfied since $A \ll 1$ (see (3.34), (3.29)). On the other hand, assumption (i) is equivalent to $\exp \left(\frac{-4 \int w^{2} d y}{A^{2} \pi}\right) \gg \varepsilon^{2}$, or $A^{2} \gg O\left(\frac{1}{\log \varepsilon^{-1}}\right)$. In summary, we have:

Proposition 3.2 Suppose that

$$
\frac{1}{\log 1 / \varepsilon} \ll A^{2} \ll 1
$$

Then the spike solution from Proposition 2.1 undergoes a Hopf bifurcation as $\tau$ is increased past $\tau=\tau_{h, \text { large }}$ where

$$
\begin{equation*}
\tau_{h, \text { large }} \sim \frac{1}{A^{2} \varepsilon^{2}} a_{0} \exp \left(\frac{-a_{1}}{A^{2}}\right) \tag{3.35}
\end{equation*}
$$

and

$$
a_{0}=\frac{4 e^{2 \gamma} \int_{0}^{\infty} w^{2}(\rho) \rho d \rho}{\pi} \approx 19.929, \quad \quad a_{1}=8 \int_{0}^{\infty} w^{2}(\rho) \rho d \rho \approx 39.474
$$

## 4. SMALL EIGENVALUES

We study the Hopf bifurcation in the small eigenvalue problem corresponding to the mode $m=1$ in (3.16). A posteriori analysis reveals that the relevant scaling is

$$
\lambda=\lambda_{0} \varepsilon^{2} \quad \tau=\tau_{0} \varepsilon^{-2}
$$

where $\lambda_{0}$ and $\tau_{0}$ are $O(1)$ with respect to $\varepsilon$. The leading order eigenvalue problem is

$$
\begin{align*}
& 0=\Phi_{0}^{\prime \prime}+\frac{1}{\rho} \Phi_{0}^{\prime}-\frac{1}{\rho^{2}} \Phi_{0}-\Phi_{0}+V_{0}^{2} \Psi_{0}+2 U_{0} V_{0} \Phi_{0}  \tag{4.36a}\\
& 0=\Psi_{0}^{\prime \prime}+\frac{1}{\rho} \Psi_{0}^{\prime}-\frac{1}{\rho^{2}} \Psi_{0}-\left(V_{0}^{2} \Psi_{0}+2 U_{0} V_{0} \Phi_{0}\right) \tag{4.36b}
\end{align*}
$$

where $\rho=|y|=|x| / \varepsilon$. The solution to (4.36) is given by:

$$
\begin{equation*}
\Phi_{0}=\frac{C}{S} V_{0 \rho} \quad \Psi_{0}=\frac{C}{S} U_{0 \rho} \tag{4.37}
\end{equation*}
$$

and satisfies the far field condition given by

$$
\begin{equation*}
\Phi_{0} \rightarrow 0, \quad \Psi_{0} \sim \frac{C}{\rho}, \quad \text { as } \quad \rho \rightarrow \infty \tag{4.38}
\end{equation*}
$$

The constant $C$ will be obtained through matching to the outer solution. The outer problem for $\psi$ is

$$
\tau_{0} \lambda \psi=\Delta \psi, \quad r \neq 0
$$

subject to $\frac{\partial \psi}{\partial r}(1)=0$ and $\psi \sim \frac{C \varepsilon e^{i \theta}}{r}$ as $r \rightarrow 0$. This yields an explicit solution

$$
\begin{equation*}
\psi=C \varepsilon \sqrt{\tau_{0} \lambda_{0}}\left(-\frac{K_{1}^{\prime}\left(\sqrt{\tau_{0} \lambda_{0}}\right)}{I_{1}^{\prime}\left(\sqrt{\tau_{0} \lambda_{0}}\right)} I_{1}\left(\sqrt{\tau_{0} \lambda_{0}} r\right)+K_{1}\left(\sqrt{\tau_{0} \lambda_{0}} r\right)\right) e^{i \theta} \tag{4.39}
\end{equation*}
$$

Recall the small-argument expansion for $K_{1}$ and $I_{1}$ is given by

$$
\begin{align*}
K_{1}(z) & \sim \frac{1}{z}+\frac{1}{2} z\left(\log z+b_{0}\right)+O\left(z^{2} \ln z\right), \quad \text { where } b_{0}=\gamma-\frac{1}{2}-\ln 2  \tag{4.40a}\\
I_{1}(z) & \sim \frac{1}{2} z+O\left(z^{3}\right) \tag{4.40b}
\end{align*}
$$

Writing (4.39) in inner variables $r=\rho \varepsilon$ and using expansions (4.40) we then obtain

$$
\begin{aligned}
\psi & \sim\left(\frac{C}{\rho}+\varepsilon^{2} \frac{1}{2} C \tau_{0} \lambda_{0} \rho\left\{\log \left(\sqrt{\tau_{0} \lambda_{0}} \rho \varepsilon\right)-\frac{K_{1}^{\prime}\left(\sqrt{\tau_{0} \lambda_{0}}\right)}{I_{1}^{\prime}\left(\sqrt{\tau_{0} \lambda_{0}}\right)}+b_{0}\right\}\right) e^{i \theta} \\
& \sim e^{i \theta}\left(\Psi_{0}(\rho)+\varepsilon \Psi_{1}(\rho)\right)
\end{aligned}
$$

The $O(1)$ terms yields the far-field behaviour for $\Psi(\rho)$ given by (4.38). The $O\left(\varepsilon^{2}\right)$ terms yield the far-field behaviour for $\Psi_{1}(\rho)$,

$$
\begin{equation*}
\Psi_{1} \sim \frac{1}{2} C \tau_{0} \lambda_{0} \rho\left\{\log \rho+\log \left(\sqrt{\tau_{0} \lambda_{0}} \varepsilon\right)-\frac{K_{1}^{\prime}\left(\sqrt{\tau_{0} \lambda_{0}}\right)}{I_{1}^{\prime}\left(\sqrt{\tau_{0} \lambda_{0}}\right)}+b_{0}\right\}, \quad \rho \gg 1 \tag{4.41}
\end{equation*}
$$

To determine $\lambda_{0}$ requires an expansion at the next order. The steady state satisfies

$$
\begin{array}{r}
\Delta V_{1}-V_{1}+2 U_{0} V_{0} V_{1}+U_{1} V_{0}^{2}=0 \\
\Delta U_{1}+A-U_{1} V_{0}^{2}-2 U_{0} V_{0} V_{1}=0 \tag{4.42b}
\end{array}
$$

and the corresponding eigenvalue problem is

$$
\begin{align*}
\lambda_{0} \Phi_{0} & =\Delta_{1} \Phi_{1}-\Phi_{1}+V_{0}^{2} \Psi_{1}+2 U_{0} V_{0} \Phi_{1}+2\left(V_{0} U_{1}+U_{0} V_{1}\right) \Phi_{0}+2 V_{0} V_{1} \Psi_{0}  \tag{4.43a}\\
\tau_{0} \lambda_{0} \Psi_{0} & =\Delta_{1} \Psi_{1}-\left(V_{0}^{2} \Psi_{1}+2 U_{0} V_{0} \Phi_{1}\right)-2\left(V_{0} U_{1}+U_{0} V_{1}\right) \Phi_{0}-2 V_{0} V_{1} \Psi_{0} \tag{4.43b}
\end{align*}
$$

subject to the far field condition (4.41).
We express (4.43) in matrix form as

$$
\begin{align*}
& \Delta_{1} \mathbf{W}+M \mathbf{W}=E \mathbf{f}_{1}+\mathbf{f}_{2}, \quad 0<\rho<\infty  \tag{4.44a}\\
& \mathbf{W} \sim\binom{0,}{C_{1} \rho \ln |\rho|+C_{2} \rho}, \quad \text { as } \rho \rightarrow \infty \tag{4.44b}
\end{align*}
$$

where:

$$
\begin{align*}
M & =\left(\begin{array}{cc}
-1+2 U_{0} V_{0} & V_{0}^{2} \\
-2 U_{0} V_{0} & -V_{0}^{2}
\end{array}\right), \quad E=\left(\begin{array}{cc}
-2\left(U_{0} V_{1}+U_{1} V_{0}\right) & -2 V_{0} V_{1} \\
2\left(U_{0} V_{1}+U_{1} V_{0}\right) & 2 V_{0} V_{1}
\end{array}\right),  \tag{4.44c}\\
\mathbf{W} & =\binom{\Phi_{1}}{\Psi_{1}}, \mathbf{f}_{1}=\binom{\Phi_{0}}{\Psi_{0}}, \quad \mathbf{f}_{2}=\binom{\lambda_{0} \Phi_{0}}{\tau_{0} \lambda_{0} \Psi_{0}},  \tag{4.44~d}\\
C_{1} & =\frac{1}{2} C \tau_{0} \lambda_{0}, \quad C_{2}=\frac{1}{2} C \tau_{0} \lambda_{0}\left\{\log \left(\sqrt{\tau_{0} \lambda_{0}} \varepsilon\right)-\frac{K_{1}^{\prime}\left(\sqrt{\tau_{0} \lambda_{0}}\right)}{I_{1}^{\prime}\left(\sqrt{\tau_{0} \lambda_{0}}\right)}+b_{0}\right\} . \tag{4.44e}
\end{align*}
$$

Let $\mathbf{P}$ be the solution of the adjoint problem,

$$
\begin{equation*}
\Delta_{1} \mathbf{P}+M^{t} \mathbf{P}=0 \tag{4.45a}
\end{equation*}
$$

subjected to the far-field behaviour condition

$$
\begin{equation*}
\mathbf{P} \sim\binom{0}{\frac{1}{\rho}} \quad \text { for } \rho \gg 1 \tag{4.45b}
\end{equation*}
$$

We multiply (4.44a) by $\rho \mathbf{P}^{t}$ and integrate to obtain

$$
\begin{equation*}
\int_{0}^{R} \mathbf{P}^{t}\left(\Delta_{1} \mathbf{W}+M \cdot \mathbf{W}\right) \rho d \rho=\int_{0}^{R} \mathbf{P}^{t} \cdot\left(E \mathbf{f}_{1}+\mathbf{f}_{2}\right) \rho d \rho \tag{4.46}
\end{equation*}
$$

Here, $R$ is a big number which we will take to infinity later. Integrating by parts, the left hand side becomes

$$
\begin{align*}
\int_{0}^{R} \mathbf{P}^{t}\left(\Delta_{1} \mathbf{W}+M \cdot \mathbf{W}\right) \rho d \rho & =\left(\mathbf{P}^{t} \cdot\left(\rho \frac{\partial \mathbf{W}}{\partial \rho}\right)-\left(\rho \frac{\partial \mathbf{P}^{t}}{\partial \rho}\right) \cdot \mathbf{W}\right)_{\rho=R}  \tag{4.47}\\
& =\left(2 C_{1} \ln R+2 C_{2}+C_{1}\right) \tag{4.48}
\end{align*}
$$

To calculate the right hand side of (4.46), we introduce $\mathbf{N}=\left(\frac{\partial V_{1}}{\partial \rho}, \frac{\partial U_{1}}{\partial \rho}\right)^{t}$. Upon differentiating the system for $V_{1}$ and $U_{1}$ with respect to $\rho$, we obtain

$$
\Delta_{1} \mathbf{N}+M \cdot \mathbf{N}=\binom{-2\left(U_{0} V_{0}\right)_{\rho} V_{1}-\left(V_{0}^{2}\right)_{\rho} U_{1}}{2\left(U_{0} V_{0}\right)_{\rho} V_{1}+\left(V_{0}^{2}\right)_{\rho} U_{1}}
$$

The key observation is that

$$
E \mathbf{f}_{1}=\frac{C}{S}\binom{-2\left(U_{0} V_{0}\right)_{\rho} V_{1}-\left(V_{0}^{2}\right)_{\rho} U_{1}}{2\left(U_{0} V_{0}\right)_{\rho} V_{1}+\left(V_{0}^{2}\right)_{\rho} U_{1}}=\frac{C}{S}\left(\Delta_{1} \mathbf{N}+M \cdot \mathbf{N}\right)
$$

It follows that

$$
\begin{aligned}
\int_{0}^{R} \mathbf{P}^{t} \cdot\left(E \cdot \mathbf{f}_{1}\right) \rho d \rho & =\frac{C}{S} \int_{0}^{R} \mathbf{P}^{t} \cdot\left(\Delta_{1} \mathbf{N}+M \cdot \mathbf{N}\right) \rho d \rho \\
& =\left.\frac{C}{S}\left(\mathbf{P}^{t} \cdot\left(\rho \frac{\partial \mathbf{N}}{\partial \rho}\right)-\left(\rho \frac{\partial \mathbf{P}^{t}}{\partial \rho}\right) \cdot \mathbf{N}\right)\right|_{\rho=R} \\
& =-2 C
\end{aligned}
$$

Next we simplify

$$
\int_{0}^{R} \mathbf{P}^{t} \cdot \mathbf{f}_{2} \rho d \rho=\int_{0}^{R}\left(P_{1} \Phi_{0}+\tau_{0} P_{2} \Psi_{0}\right) \lambda_{0} \rho d \rho
$$

and we further compute

$$
\begin{aligned}
\int_{0}^{R} \tau_{0} P_{2} \Psi_{0} \lambda_{0} \rho d \rho & =\left.\frac{C}{S} \tau_{0} P_{2} U_{0} \lambda_{0} \rho\right|_{0} ^{R}-\frac{C}{S} \int_{0}^{R} \tau_{0}\left(P_{2} \rho\right)_{\rho} U_{0} \lambda_{0} d \rho \\
& =\frac{C}{S} \tau_{0} \lambda_{0}(S \log (R)+\chi(S))-\frac{C}{S} \int_{0}^{R} \tau_{0}\left(P_{2} \rho\right)_{\rho} U_{0} \lambda_{0} d \rho
\end{aligned}
$$

In summary, we obtain that the right hand side of (4.46) simplifies to

$$
\begin{equation*}
\int_{0}^{R} \mathbf{P}^{t} \cdot\left(E \mathbf{f}_{1}+\mathbf{f}_{\mathbf{2}}\right) \rho d \rho=\lambda_{0} \int_{0}^{R} P_{1} \Phi_{0} \rho d \rho+C \tau_{0} \lambda_{0} \log (R)+\frac{C \tau_{0} \lambda_{0}}{S} \chi(S)-\frac{C}{S} \int_{0}^{R} \tau_{0}\left(P_{2} \rho\right)_{\rho} U_{0} \lambda_{0} d \rho-2 C \tag{4.49}
\end{equation*}
$$

Equating (4.48) and (4.49), note that the $\log R$ terms cancel each other out and after factoring out $C$, and we finally obtain

$$
\lambda_{0} \int_{0}^{R} P_{1} \frac{1}{S} V_{0 \rho} \rho d \rho+\frac{\tau_{0} \lambda_{0}}{S} \chi(S)-\int_{0}^{R} \tau_{0}\left(P_{2} \rho\right)_{\rho} U_{0} \lambda_{0} d \rho-2=\tau_{0} \lambda_{0}\left(\log \left(\frac{\sqrt{\tau_{0} \lambda_{0}} \varepsilon}{2}\right)-\frac{K_{1}^{\prime}\left(\sqrt{\tau_{0} \lambda_{0}}\right)}{I_{1}^{\prime}\left(\sqrt{\tau_{0} \lambda_{0}}\right)}+\gamma\right)
$$

or

$$
\begin{equation*}
\frac{\lambda_{0} \kappa_{1}-\tau_{0} \lambda_{0} \kappa_{2}}{S}=\tau_{0} \lambda_{0}\left(\log \left(\frac{e^{\gamma}}{2} \sqrt{\tau_{0} \lambda_{0}} \varepsilon\right)-\frac{K_{1}^{\prime}\left(\sqrt{\tau_{0} \lambda_{0}}\right)}{I_{1}^{\prime}\left(\sqrt{\tau_{0} \lambda_{0}}\right)}\right)+2 \tag{4.50a}
\end{equation*}
$$

where $\kappa_{1}$ and $\kappa_{2}$ are given by

$$
\begin{equation*}
\kappa_{1}:=\int_{0}^{\infty} P_{1} V_{0 \rho} \rho d \rho, \quad \kappa_{2}:=\int_{0}^{\infty}\left(P_{2} \rho\right)_{\rho}\left[U_{0}-\chi(S)\right] d \rho \tag{4.50b}
\end{equation*}
$$

Next, we seek a Hopf bifurcation for (4.50). Setting $\lambda_{0}=i \lambda_{I}$ in (4.50a) and equating real and imaginary parts yields $\tau_{0} \lambda_{I}=\omega_{c}$ where $\omega_{c}$ satisfies

$$
\begin{equation*}
\omega_{c} \operatorname{Im}\left(\log \left(\sqrt{i \omega_{c}}\right)-\frac{K_{1}^{\prime}\left(\sqrt{i \omega_{c}}\right)}{I_{1}^{\prime}\left(\sqrt{i \omega_{c}}\right)}\right)-2=0 \tag{4.51a}
\end{equation*}
$$

and

$$
\begin{equation*}
\tau_{0}=\frac{\kappa_{1}}{S \operatorname{Re}\left(-\frac{K_{1}^{\prime}\left(\sqrt{i \omega_{c}}\right)}{I_{1}^{\prime}\left(\sqrt{i \omega_{c}}\right)}+\log \left(\frac{e^{\gamma}}{2} \sqrt{i \omega_{c}} \varepsilon\right)\right)+\kappa_{2}} \tag{4.51b}
\end{equation*}
$$

A remarkable fact is that the equation (4.51a) is independent of any parameters. Numerical plotting shows that there is a unique solution to (4.51a) given by

$$
\begin{equation*}
\omega_{c} \approx 3.02603687 \tag{4.51c}
\end{equation*}
$$

Expression (4.51b) is further simplifed by rewriting

$$
\operatorname{Re}\left(-\frac{K_{1}^{\prime}\left(\sqrt{i \omega_{c}}\right)}{I_{1}^{\prime}\left(\sqrt{i \omega_{c}}\right)}+\log \left(\frac{e^{\gamma}}{2} \sqrt{i \omega_{c}} \varepsilon\right)\right)=\alpha_{1}+\log (\varepsilon)
$$

where

$$
\begin{equation*}
\alpha_{1}:=\operatorname{Re}\left(-\frac{K_{1}^{\prime}\left(\sqrt{i \omega_{c}}\right)}{I_{1}^{\prime}\left(\sqrt{i \omega_{c}}\right)}\right)+\log \left(\frac{e^{\gamma}}{2} \sqrt{\omega_{c}}\right)=-0.14623425 \tag{4.52}
\end{equation*}
$$

In general, the constants $\kappa_{1}, \kappa_{2}$ must be computed numerically. However asymptotic expansion is available in the intermediate regime, when $A$ is small, and is given in Appendix A. We summarize.
Proposition 4.1 The translational eigenvalue corrsponding to mode $m=1$ of the steady state in Proposition 2.1 undergoes a Hopf bifurcation as $\tau$ is increased past $\tau \sim \tau_{h, s m a l l}$ where

$$
\begin{equation*}
\tau_{h, \text { small }}=\frac{1}{\varepsilon^{2}} \frac{\kappa_{1}}{\frac{A}{2}\left(\log \varepsilon+\alpha_{1}\right)+\kappa_{2}} \tag{4.53}
\end{equation*}
$$

The constants $\kappa_{i}$ are independent of $\varepsilon$ (depend only on $A$ ) and are given in (4.50b). The constant $\alpha_{1}=-0.1462342$ is a universal constant defined through (4.52). In the asymptotic limit $A \ll 1$, the formula (4.53) simplifies to

$$
\begin{equation*}
\tau_{h, \text { small }}=\frac{1}{\varepsilon^{2} A^{2}} \frac{2 \kappa_{10}}{\log \frac{1}{\varepsilon}-\alpha_{1}-2 \kappa_{20}} \tag{4.54}
\end{equation*}
$$

where

$$
\kappa_{10} \approx 9.86855 ; \quad \kappa_{20} \approx 0.1441
$$

whose exact value is derived in Appendix $A$.
Figure 2 shows a very good agreement between the full numerical simulations of the eigenvalue problem (3.14) and formula (4.54).

## 5. THRESHOLD CROSSING

As Figure 2 shows, the Hopf curves $\tau_{h, l a r g e}$ and $\tau_{h, s m a l l}$ intersect as $A$ is increased at some critical value $A=A_{c}$. Having computed asymptotically the Hopf bifurcations for both small and large eignevalues, we are finally in position to determine this crossing by equating $\tau_{h, \text { large }}=\tau_{h, s m a l l}$ (where $\tau_{h, l a r g e}$ and $\tau_{h, s m a l l}$ are given in Propositions 3.2 and 4.1, respectively). Solving for $A$ yields

$$
\begin{equation*}
A_{c} \sim \frac{c_{1}}{\left[\ln \left(c_{2} \ln \frac{1}{\varepsilon}+c_{3}\right)\right]^{1 / 2}} \tag{5.55}
\end{equation*}
$$

where

$$
\begin{aligned}
& c_{1}=a_{1}^{1 / 2} \approx 6.2828 \\
& c_{2}=\frac{e^{2 \gamma}}{\pi} \approx 1.00975 \\
& c_{3}=\frac{\left(-\alpha_{1}-2 \kappa_{20}\right) a_{0}}{2 \kappa_{10}} \approx-0.14334
\end{aligned}
$$

From the formulas for $\tau_{h, \text { large }}$ and $\tau_{h, s m a l l}$, it is clear that if $A<A_{c}$, the height oscillations are triggered before position oscillations, whereas the opposite is true if $A>A_{c}$. This is the main result of the paper. We summarize.

Proposition 5.1 Let $A_{c}$ as given in (1.3) with $\varepsilon \ll 1$. Suppose that $A<A_{c}$. Then height oscillations are triggered before the position oscillations as $\tau$ is increased just past $\tau_{h, l a r g e}$. Suppose that $A>A_{c}$. Then position oscillations are triggered before height oscillations as $\tau$ is increased just past $\tau_{h, s m a l l}$.

Note that the derivation required that $O\left(\frac{1}{\log \frac{1}{\varepsilon}}\right) \ll A^{2} \ll O(1)$. Both of these conditions are clearly satisfied in the critical regime $A^{2}=O\left(A_{c}^{2}\right)=O\left(\frac{1}{\log \left(\log \frac{1}{\varepsilon}\right)}\right)$.

Although in theory, the formula for $A_{c}$ is is valid as $\varepsilon \rightarrow 0$, the $\log$-log scaling has a horrible convergence rate. It is then all the more surprising that the formula (1.3) is able to predict the threshold within a reasonable accuracy, even when $\varepsilon=0.02$. To further validate this result, we computed $A_{c}$ numerically up to $\varepsilon=O\left(10^{-3}\right)$. The result is summarized in the table in Figure 2. Attempting to compute at such small $\varepsilon$ values required the use of a non-uniform grid to compute eigenvalues numerically. We then used a numerical root solver and continuation to adjust $A$ until $\tau_{h, \text { large }}=\tau_{h, \text { small }}$. We validated our computations by doubling the number of meshpoints. The prediction given by (1.3) is increasingly accurate with each halfing of $\varepsilon$, although as expected from a log-log scaling, the improvement in accuracy is very slow.

## 6. ROTATING SPOT

When the spike is destabilized via translational instabilities, it starts to move as illustrated in figure 1, and may eventually settling into a circular orbit, rotating with some frequency $\omega_{0}$ around some radius $r_{0}$. The goal of this section is to compute $\omega_{0}$ and $r_{0}$ asymptotically. Before proceeding, it is convenient to rescale

$$
\tau=\frac{\tau_{0}}{\varepsilon^{2}}, \quad s=\varepsilon^{2} t
$$

so that the problem (1.1) becomes

$$
\begin{equation*}
\varepsilon^{2} v_{s}=\varepsilon^{2} \Delta v-v+v^{2} u, \quad \tau_{0} u_{s}=\Delta u+A-\frac{v^{2} u}{\varepsilon^{2}} \tag{6.56}
\end{equation*}
$$

Let $x_{0}(s)$ be the location of the spot. To make further progress, we make the anzatz that the spot travels along a circle of radius $r_{0}$ with constant angular velocity $\omega_{0}$, so that both $u$ and $v$ undergo a rigid rotation. That is, we assume

$$
\begin{equation*}
x_{0}(s)=e^{i \omega_{0} s} r_{0} \tag{6.57}
\end{equation*}
$$

and $u(x, s)=u\left(x e^{-i \omega_{0} s}\right), v(x, s)=v\left(x e^{-i \omega_{0} s}\right)$. We will estimate inner and outer region and perform matching in order to obtain a solvability condition which will determine the radius and the angular velocity of the spot.


FIG. 4. Radius of rotating spot as a function of $\tau_{0}=\tau \varepsilon^{2}$. Parameter values are $A=8$ and $\varepsilon=0.02$. Dashed line is the asymptotic theory given by Proposition 6.1. Circle is the average long-time radius as observed from direct numerical simulations. Inserts show the long-time spike trajectory for $\tau_{0}$ as indicated (with numerical trajectory shown in solid line and the theoretical rotating-spot trajectory of radius $r_{0}$ shown in dashed line).

Outer region. Away from spike location, we estimate the outer problem for $u$ by

$$
\begin{equation*}
\Delta u+A=2 \pi S \delta\left(x-x_{0}\right)+\tau_{0} u_{s} \tag{6.58}
\end{equation*}
$$

with Neumann boundary condition $\partial_{n} u=0, \quad x \in \partial \Omega$. Here, $S$ is defined by

$$
2 \pi S=\int \frac{u v^{2}}{\varepsilon^{2}} d x \sim \int_{\mathbb{R}^{2}} U V^{2} d y, \quad y=\frac{x-x_{0}(s)}{\varepsilon}
$$

The relation between $S$ and $A$ is determined by integrating (6.58) to obtain

$$
\begin{equation*}
\pi A=2 \pi S+\tau_{0} \frac{d}{d s}\left(\int_{\Omega} u d x\right) . \tag{6.59}
\end{equation*}
$$

But since we assumed that $u$ is rigidly rotating, the integral term $\int_{\Omega} u d x$ is independent of time $s$ so that - just as for stationary spot - (6.59) simplifies to

$$
\begin{equation*}
S=\frac{A}{2} \tag{6.60}
\end{equation*}
$$

We write $u$ as

$$
u(x, s)=2 S G(x, s)+C
$$

where $G$ satisfies

$$
\left\{\begin{array}{c}
\Delta G+1=\pi \delta\left(x-x_{0}(s)\right)+\tau_{0} G_{s}  \tag{6.61}\\
\partial_{n} G=0, \quad x \in \partial \Omega ; \quad \int G d x=0
\end{array} ; \quad x_{0}(s)=e^{i \omega_{0} s} r_{0},\right.
$$

and $C$ is some (irrelevant for dynamics) constant. We now show that to leading order, (6.61) has the following singularity structure

$$
\begin{equation*}
G(x)=\frac{1}{2} \log \left|x-x_{0}\right|+\frac{\tau_{0}}{4} \dot{x}_{0} \cdot\left(x-x_{0}\right) \log \left|x-x_{0}\right|+R_{0}\left(x, x_{0}\right) . \tag{6.62}
\end{equation*}
$$

where $R_{0}$ is the "regular part", in the sense that its gradient exists at $x=x_{0}$; we give explicit expression for $R_{0}$ in appendix C. Note that when $\tau_{0}=0$, this corresponds to the usual Modified Green's function on a disk; however the non-zero $\tau_{0}$ induces an an additional singularity term $\dot{x}_{0} \cdot\left(x-x_{0}\right) \log \left|x-x_{0}\right|$. This latter term is "singular" in the sense that its gradient is infinite as $x \rightarrow x_{0}$ and therefore needs to be "peeled off".

To see where this singularity comes from, first consider the source that moves along $y$-axis, with some speed $c$, $x_{0}(s)=(0, c s)$ on all of space; the free-space moving source Green's function then satisfies

$$
\begin{equation*}
\Delta G=\tau_{0} G_{s}+\pi \delta\left(x-x_{0}(s)\right) ; \quad x \in \mathbb{R}^{2} \text { and } x_{0}(s)=(0, c s) \tag{6.63}
\end{equation*}
$$

In this case, transforming into co-moving coordinates $x=(\xi, \eta)+(0, c s)$ yields

$$
G_{\xi \xi}+G_{\eta \eta}+c \tau_{0} G_{\eta}=\pi \delta(\xi) \delta(\eta)
$$

This problem has an exact solution of the form

$$
\begin{equation*}
G(\xi, \eta)=-\frac{1}{2} e^{\frac{-c \tau_{0}}{2} \eta} K_{0}\left(\frac{c \tau_{0}}{2} r\right), \quad r=\sqrt{\xi^{2}+\eta^{2}} \tag{6.64}
\end{equation*}
$$

We then expand for small $r$ and $y$ using Taylor expansions $K_{0}(z) \sim-\log z, e^{c / 2 \eta} \sim 1+\frac{c \tau_{0}}{2} \eta$ which yields

$$
\begin{equation*}
G(\xi, \eta) \sim \frac{1}{2}\left(1-\frac{c \tau_{0}}{2} \eta\right) \log r+\ldots \tag{6.65}
\end{equation*}
$$

This also explains the choice of the constant $\frac{-1}{2}$ in (6.64) which gives the correct leading order behaviour $G \sim \frac{1}{2} \log r$ independent of $c \tau_{0}$. Replacing $c \eta$ by $\dot{x}_{0} \cdot\left(x-x_{0}\right)$ and $r$ by $\left|x-x_{0}\right|$ indeed yields the singularity structure (6.62). Further expanding $x$ near $x_{0}$, the outer problem for $u(x)$ is then given by

$$
\begin{equation*}
u(x) \sim S \log \left|x-x_{0}\right|-\frac{\tau_{0} S}{2} \dot{x}_{0} \cdot\left(x-x_{0}\right) \log \left|x-x_{0}\right|+2 S\left(x-x_{0}\right) \cdot \nabla R_{0}+C, \quad x \rightarrow x_{0} \tag{6.66}
\end{equation*}
$$

where $\nabla R_{0}=\left.\nabla_{x} R_{0}\left(x, x_{0}\right)\right|_{x=x_{0}}$ and $C$ is some constant.
Inner region. In the inner region near the spot, we rescale

$$
y=\frac{x-x_{0}(s)}{\varepsilon} ; \quad v(x, t)=V(y) \quad u(x, t)=U(y)
$$

Then $V, U$ satisfies

$$
\left\{\begin{align*}
-\varepsilon \nabla_{y} V \frac{d x_{0}}{d s} & =\Delta_{y} V-V+U V^{2}  \tag{6.67}\\
-\tau_{0} \varepsilon \nabla_{y} U \frac{d x_{0}}{d s} & =\Delta_{y} U+A \varepsilon^{2}-U V^{2}
\end{align*}\right.
$$

We then expand in $\varepsilon$,

$$
\begin{equation*}
U=U_{0}+\varepsilon U_{1}+\cdots, \quad V=V_{0}+\varepsilon V_{1}+\cdots \tag{6.68}
\end{equation*}
$$

At the leading order we have

$$
\left\{\begin{array}{c}
\Delta V_{0}-V_{0}+U_{0} V_{0}^{2}=0  \tag{6.69}\\
\Delta U_{0}-U_{0} V_{0}^{2}=0
\end{array}\right.
$$

At the next order we obtain

$$
\left\{\begin{array}{c}
\Delta V_{1}-V_{1}+2 U_{0} V_{0} V_{1}+V_{0}^{2} U_{1}=-\nabla V_{0} \cdot \dot{x}_{0}  \tag{6.70}\\
\Delta U_{1}-2 U_{0} V_{0} V_{1}-V_{0}^{2} U_{1}=-\tau_{0} \nabla U_{0} \cdot \dot{x}_{0}
\end{array}\right.
$$

We assume that $V_{0}, V_{1}$ decays exponentially in the far field $|y| \gg 1$. To obtain the far-field behaviour for $U_{0}$ and $U_{1}$, we rewrite the outer expansion (6.66) in the inner variables. This yields

$$
\begin{equation*}
u(x)=S \log \varepsilon|y|+\frac{\tau_{0} S}{2} \dot{x}_{0} \cdot y \varepsilon \log \varepsilon|y|+2 S \nabla R_{0} \cdot y \varepsilon+C \tag{6.71}
\end{equation*}
$$

Upon collecting like terms in $\varepsilon$ (while treating $\log \varepsilon$ as an $O(1)$ constant with respect to $\varepsilon$ ), we obtain

$$
\begin{equation*}
U_{0} \sim S \log |y|+\chi(S), \quad|y| \gg 1 \tag{6.72}
\end{equation*}
$$

The function $\chi(S)$ is the same as in (2.7) and the constant $C$ in (6.71) determined through the relationship $\chi(S)=$ $S \log \varepsilon+C$.

At the next order we obtain

$$
U_{1} \sim \frac{\tau_{0} S}{2} \dot{x}_{0} \cdot y \log |y|+\left(2 S \nabla R_{0}-\frac{\tau_{0} S}{2} \dot{x}_{0} \log \varepsilon\right) \cdot y, \quad|y| \gg 1
$$

Following the derivation in $\S 4$, we rewrite the system (6.70) as

$$
\begin{align*}
& \Delta W+M \cdot W=f \quad, \quad y \in \mathbb{R}^{2}  \tag{6.73a}\\
& W \sim\left(0,-\frac{S \tau_{0}}{2} \dot{x}_{0} \cdot y \ln |y|+\vec{b} \cdot y\right)^{t}, \quad \text { as }|y| \rightarrow \infty \tag{6.73b}
\end{align*}
$$

where

$$
\begin{gather*}
\vec{b}=\frac{\tau_{0} S}{2} \dot{x}_{0} \log \varepsilon+2 S \nabla R_{0}  \tag{6.73c}\\
M=\left(\begin{array}{cc}
-1+2 U_{0} V_{0} & V_{0}^{2} \\
-2 U_{0} V_{0} & -V_{0}^{2}
\end{array}\right), \quad \mathbf{W}=\binom{V_{1}}{U_{1}}, \quad \mathbf{f}=\binom{-\nabla_{y} V_{0} \cdot \dot{x}_{0}}{-\tau_{0} \nabla_{y} U_{0} \cdot \dot{x}_{0}} . \tag{6.73d}
\end{gather*}
$$

As in $\S 4$, to formulate the solvability condition, we let $P(\rho)=\left(P_{1}(\rho), P_{2}(\rho)\right)^{t}$ be the solution to the homogeneous adjoint problem associated with (4.44a), given by (4.45). Define

$$
\begin{equation*}
P_{c}=P(\rho) \cos \theta, \quad P_{s}=P(\rho) \sin \theta \tag{6.74}
\end{equation*}
$$

where $\cos \theta=\frac{y_{1}}{|y|}$ and $\sin \theta=\frac{y_{2}}{|y|}$; note that $P_{c}$ and $P_{s}$ both satisfy $\Delta P+M^{t} P=0$.
Multiply (6.73a) by $P_{c}^{t}$ and integrate by parts over a ball of large radius $R$ to obtian the solvablity condition

$$
\begin{equation*}
\int_{B_{R}} P_{c}^{t} \cdot f d y=\int_{\partial B_{R}} P_{c}^{t} \cdot \partial_{\rho} W-W \cdot \partial_{\rho} P_{c}^{t} d y \tag{6.75}
\end{equation*}
$$

The left hand side of (6.75) simplifies to

$$
\begin{align*}
\int_{B_{R}} P_{c}^{t} \cdot f d y & =-\pi \int_{0}^{R}\left(P_{1} V_{0 \rho}+\tau_{0} P_{2} U_{0 \rho}\right) \dot{x}_{01} \rho d \rho  \tag{6.76}\\
& \sim-\pi \dot{x}_{01} \pi\left(\kappa_{1}+\tau_{0} S \log R-\tau_{0} \kappa_{2}\right) \tag{6.77}
\end{align*}
$$

where $\kappa_{1}, \kappa_{2}$ are defined in (4.50b).
The right and side of (6.75) simplifies to

$$
\begin{equation*}
\int_{\partial B_{R}} P_{c}^{t} \cdot \partial_{\rho} W-W \cdot \partial_{\rho} P_{c}^{t} d y=\pi\left(-S \tau_{0} \dot{x}_{01}\left[\frac{1}{2}+\ln R\right]+2 b_{1}\right) \tag{6.78}
\end{equation*}
$$

where $b_{1}$ is the first component of vector $b$ in (6.73c). Equating (6.77) and (6.78), note that the $\log R$ terms cancel each other out and we finally obtain

$$
-\left(\kappa_{1}-\tau_{0} \kappa_{2}\right) \dot{x}_{01}=-\frac{S}{2} \tau_{0} \dot{x}_{01}+2 b_{1}
$$

The second solvability condition involving $\dot{x}_{01}$ is obtained similarly by using $P_{s}$ instead of $P_{c}$. The two solvability conditions together yield

$$
\begin{equation*}
-\left(\kappa_{1}-\tau_{0} \kappa_{2}\right) \dot{x}_{0}=-\frac{S}{2} \tau_{0} \dot{x}_{0}+2\left(2 S \nabla R_{0}-\frac{\tau_{0} S}{2} \dot{x}_{0} \log \varepsilon\right) \tag{6.79}
\end{equation*}
$$



FIG. 5. Complex spot trajectories. Left the domain is a unit disk. Right: the domain is a square of area $\pi$.

Solving for $\dot{x}_{0}$ and using $S=A / 2$ then yields

$$
\begin{equation*}
\frac{d}{d s} x_{0}=\beta \nabla R_{0} \quad \text { where } \beta=\frac{1}{\tau_{0}\left(\frac{1}{4} \log \varepsilon+\frac{1}{8}+\frac{\kappa_{2}}{2 A}\right)-\frac{\kappa_{1}}{2 A}} \tag{6.80}
\end{equation*}
$$

In Appendix C we derive an exact expansion for $R_{0}$ in terms of an infinite series of Bessel functions. By symmetry, we may assume without loss of generality that $x_{0}$ lies on the positive x -axis (i.e. $\left.s=0\right)$. Then $\left.\frac{d}{d s} x_{0}\right|_{s=0}=\left(0, \omega_{0} r_{0}\right)$ and we write:

$$
\begin{equation*}
\left.\nabla R_{0}\left(x, x_{0}\right)\right|_{x_{0}=\left(r_{0}, 0\right), x=x_{0}}=\left(F_{1}\left(r_{0}, \omega\right), F_{2}\left(r_{0}, \omega\right)\right), \quad \text { where } \omega=\omega_{0} \tau_{0} \tag{6.81}
\end{equation*}
$$

with $F_{1}, F_{2}$ given in (C.7). Equation (6.81) is then equivalent to $F_{1}=0, \omega_{0} r_{0}=F_{2}$, or

$$
\begin{equation*}
F_{1}\left(r_{0}, \omega\right)=0 ; \quad \tau_{0}=\frac{2 \kappa_{1}}{A\left(\log \varepsilon+\frac{1}{2}-\frac{4 F_{2}\left(r_{0}, \omega\right)}{\omega r_{0}}\right)+\kappa_{2}} ; \quad \omega_{0}=\omega / \tau_{0} \tag{6.82}
\end{equation*}
$$

In addition, as we show in Appendix C, the threshold $\tau_{h, s m a l l}$ of Proposition 4.1 is recovered in the limit $r_{0} \rightarrow 0$. We summarize our construction as follows.

Proposition 6.1 The Schnakenberg model (1.1) on a unit disk admits a rotating spot solution for $\tau>\tau_{h, s m a l l}$, where $\tau_{h, s m a l l}$ is the Hopf bifurcation value with respect to translational eigenvalues as given in Proposition 4.1. The spot center $x_{0}=r_{0} e^{i \omega_{0} \varepsilon^{2} t}$ rotates with angular velocity $\omega_{0} \varepsilon^{2}$ and radius $r_{0}$. as determined through (6.82).

Figure 4 shows a comparison between the numerical simulations of the full system (1.1) and the asymptotic prediction for the radius of the rotating spot. For example take $\tau_{0}=0.15, A=8, \varepsilon=0.02$. Then using (4.50b) we first compute $\kappa_{1}=-1.2938$ and $\kappa_{2}=3.54334$ by solving the radial core problem and the adjoint eigenvalue problem using a boundary value problem solver in Matlab (bvp4c). From (6.82) we then obtain $r_{0}=0.669$ and $\omega=6.1994$. Full numerical simulations of the original model (1.1) exhibit a rotating spot whose radius is $r_{0, \text { numeric }} \approx 0.57$, in good agreement with the theoretical prediction. Although the Proposition 6.1 applies for any $\tau>\tau_{h, s m a l l}$, the rotating spot solution is not always stable as Figure 4 shows. For example when $\tau_{0}=0.11$, the numerical solution appears to be in the shape of an ellipse whereas for $\tau_{0}=0.18$ the radius is close to the theoretical prediction but appears to vary with time, generating an annular region. More complex trajectories are possible as shown in Figure 5.

## 7. DISCUSSION

We have used formal asymptotics to compute Hopf bifurcation thresholds $\tau=\tau_{h, l a r g e}$ and $\tau=\tau_{h, s m a l l}$ that induce spike oscillations in either height $\left(\tau_{h, \text { large }}\right)$ or position $\left(\tau_{h, s m a l l}\right)$ for the Schankenberg model. These two thresholds
cross at $A=A_{c}$. That is, height oscillations dominate $\left(\tau_{h, l a r g e}<\tau_{h, s m a l l}\right)$ when $A<A_{c}$ whereas position oscillations dominate $\left(\tau_{h, \text { small }}<\tau_{h, \text { large }}\right)$ when $A>A_{c}$, where $A_{c}$ given by (1.3) has an $O(1 / \log (\log \varepsilon))$ scaling. Despite the extremely slow decay of $A_{c}$ as $\varepsilon \rightarrow 0$, the asymptotically computed value of $A_{c}$ agrees surprisingly well with numerics even with $\varepsilon=0.01$ (the constants in (1.3) are very important to get a good agreement). We remark that in one dimension, a similar "double-hopf" point was found in several papers [7-9]. However it has an algebraic scaling $A_{c}=O\left(\varepsilon^{1 / 6}\right)([8])$.

In the regime $\tau_{h, \text { small }}<\tau<\tau_{h, \text { large }}$, we have consturcted a periodic spike solution consisting of a rotating spot inside a unit disk, and computed the radius and speed of the rotation by expanding the underlying Green's function in terms of complex Bessel series. Numerical experiments suggest that the rotating solution is not always stable see Figure 4. In particular, for $\tau$ just slightly above the bifurcation point $\tau_{h, s m a l l}$, the spike trajectory is an ellipse, whereas for $\tau$ sufficiently large, the spot path fills out an annulus. It would be a very interesting to study the stability of these rotating spots.

It would be interesting to study more general spot motion and for more general domains. Figure 5 gives some idea of possible trajectories. A preliminary goal is to derive and numerically simulate the reduced equations of motion. The reduced equations of motion comprise a coupled PDE-ODE system with a moving source, analogous to the equations derived in $\S 6$. The numerical difficulty is the ODE for the source location requires an extraction of a very weakly singular part of the moving Green's function.

Circular spot motion is intimately related to the model of a small rotating trap inside an insulated unit disk, which was recently studied in $[26,27]$. There, the main goal was to minimize the mean first passage time (MFPT) for a rotating trap $x_{0}=r_{0} e^{\omega t}$ (or several rotating traps) as a function of its radius $r_{0}$ and its angular velocity $\omega$. It turned out that the optimal radius $r_{0}$ and velocity $\omega$ have precisely the same relation $F_{1}\left(r_{0}, \omega\right)=0$ as we found in Proposition 6.1. As a result, for small angular velocity $\left(\omega<\omega_{c}\right)$, it was optimal for the trap to be located at the origin, whereas for $\omega>\omega_{c}$ it was better for the spot to move. This is the precise analogue of the Hopf bifurcation computed in Proposition 4.1.

Spot motion was also observed for a three-component gas-discharge system [28]. There, the authors also analysed complex spot dynamics, including spot collision and splitting. The initial instability inducing spot motion in this system was further analysed in detail in [29] where theoretical and numerical study of the bifurcation from a stationary to a moving spot was performed. Let us also mention the work [30] where complex motion of a self-propelled deformable particle was studied.

While in many aspects, GM, GS and Schakengerg models are very similar mathematically, the oscillations of spot positions have never been observed in GM model. It would be interesting to have a better understanding of the kind of general conditions that are needed to observe position oscillations.

## Appendix A: Estimating $\kappa_{1}, \kappa_{2}$ for small $A$.

In this appendix we compute the asymptotic expansion for $\kappa_{1}, \kappa_{2}$ given by (4.50b) for small $A$. In this limit, we recall from (2.9) that $U_{0} \sim \chi=\sigma^{-1}$ and $V_{0} \sim \sigma w$ where $\sigma=\frac{A}{2 \int_{0}^{\infty} w^{2} \rho d \rho}$. The adjoint problem (4.45) simplifies to

$$
\begin{align*}
& \left(\partial_{\rho \rho}+\rho^{-1} \partial_{\rho}-\rho^{-2}\right) P_{1}-P_{1}+2 w P_{1}-2 w P_{2} \sim 0, \quad 0<\rho<\infty, \quad P_{1} \sim 0 \text { as } \rho \rightarrow \infty  \tag{A.1a}\\
& \left(\partial_{\rho \rho}+\rho^{-1} \partial_{\rho}-\rho^{-2}\right) P_{2}+\sigma^{2} w^{2}\left(P_{1}-P_{2}\right) \sim 0, \quad 0<\rho<\infty, \quad P_{2} \sim \rho^{-1}, \quad \text { as } \rho \rightarrow \infty \tag{A.1b}
\end{align*}
$$

The solution to this limiting system is given by

$$
\begin{equation*}
P_{1}=\sigma^{-2}\left(-\frac{3 w_{\rho}}{\int_{0}^{\infty} w^{3} s d s}+O\left(\sigma^{2}\right)\right), \quad P_{2}=\frac{1}{\rho} \frac{\int_{0}^{\rho} w^{3} s d s}{\int_{0}^{\infty} w^{3} s d s}+O\left(\sigma^{2}\right) \tag{A.2}
\end{equation*}
$$

and we then obtain

$$
\begin{equation*}
\kappa_{1} \sim \frac{-3 \int_{0}^{\infty} w_{\rho}^{2} \rho d \rho}{\int_{0}^{\infty} w^{3} \rho d \rho} \sigma^{-1}, \quad \kappa_{2} \sim \frac{1}{\int_{0}^{\infty} w^{3} \rho d \rho} \int_{0}^{\infty}\left[U_{0}-\sigma^{-1}\right] w^{3} \rho d \rho \tag{A.3}
\end{equation*}
$$

To compute $\kappa_{2}$ further, we let $U_{0}-\sigma^{-1}=\sigma \hat{U}+O\left(\sigma^{2}\right)$ where $\hat{U}$ satisfies

$$
\Delta \hat{U}=w^{2}
$$

subject to the far-field condition $\hat{U} \sim S \log \rho+0, \rho \gg 1$. The solution to $\hat{U}$ is given by

$$
\hat{U}(\rho)=\int_{0}^{\rho} \frac{F(s)}{s} d s-\int_{1}^{\infty} \frac{F(s)-F(\infty)}{s} d s-\int_{0}^{1} \frac{F(s)}{s} d s
$$

where

$$
F(s)=\int_{0}^{s} w^{2}(\rho) \rho d \rho
$$

$\kappa_{2}$ is then given by

$$
\kappa_{2} \sim \sigma \frac{\int_{0}^{\infty} \hat{U} w^{3} \rho d \rho}{\int_{0}^{\infty} w^{3} \rho d \rho}
$$

The integrals $\int_{0}^{\infty} \hat{U} w^{3} \rho d \rho$ is computed numerically. In summary, we obtain the following expansions for $\kappa_{1}, \kappa_{2}$ :

$$
\kappa_{1} \sim-\frac{1}{A} \kappa_{10} ; \quad \kappa_{2} \sim A \kappa_{20}
$$

where (using (B.4)),

$$
\kappa_{10}=2 \int_{0}^{\infty} w^{2} \rho d \rho ; \quad \kappa_{20}=\frac{\int_{0}^{\infty} \hat{U} w^{3} \rho d \rho}{2 \int_{0}^{\infty} w^{2} \rho d \rho \int_{0}^{\infty} w^{3} \rho d \rho}
$$

The numerical estimates for $\kappa_{1}$ and $\kappa_{2}$, computed using numerical quadrature, are

$$
\kappa_{10} \approx 9.8686 ; \quad \kappa_{20} \approx 0.1441
$$

## Appendix B: Some properties of function $w$

We start from the ground state $w(y)=w(\rho), \quad \rho=|y|$. It satisfies

$$
\begin{equation*}
w^{\prime \prime}+\frac{w^{\prime}}{\rho}-w+w^{2}=0, \quad w^{\prime}(0)=0, \quad w \rightarrow 0 \quad \text { as } \rho \rightarrow \infty \tag{B.1}
\end{equation*}
$$

Multipling (B.1) by $w \rho$ and $w^{\prime} \rho^{2}$ respectively and integrating over the domain yields

$$
\begin{array}{r}
-\int_{0}^{\infty} w_{\rho}^{2} \rho d \rho-\int_{0}^{\infty} w^{2} \rho d \rho+\int_{0}^{\infty} w^{3} \rho d \rho=0 \\
\int_{0}^{\infty} w^{2} \rho d \rho-\frac{2}{3} \int_{0}^{\infty} w^{3} \rho d \rho=0 \tag{B.3}
\end{array}
$$

Combining these two equations leads to

$$
\begin{equation*}
\frac{\int_{0}^{\infty} w_{\rho}^{2} \rho d \rho}{\int_{0}^{\infty} w^{3} \rho d \rho}=\frac{1}{3}, \quad \frac{\int_{0}^{\infty} w^{2} \rho d \rho}{\int_{0}^{\infty} w^{3} \rho d \rho}=\frac{2}{3} \tag{B.4}
\end{equation*}
$$

Finally, we will use the following numerical estimate:

$$
\int_{0}^{\infty} w^{2} \rho d \rho \approx 4.9343
$$

It is obtained by solving (B.1) using Matlab's boundary value problem solver bvp4c, then using numerical quadrature for the resulting integral.

## Appendix C: Green's function for rotating spot

In this appendix we compute explicitly gradient of the regular part of the rotating Green's function, defined through (6.61), (6.62). In the rotating frame, the Green's function $G$ from (6.61) satisfies

$$
\begin{align*}
& \frac{\partial^{2} G}{\partial r^{2}}+\frac{1}{r} \frac{\partial G}{\partial r}+\frac{1}{r^{2}} \frac{\partial^{2} G}{\partial \theta^{2}}+\omega \frac{\partial G}{\partial \theta}=1-\pi \delta\left(r-r_{0}\right) \delta(\theta)  \tag{C.1}\\
& \partial_{r} G=0 \quad \mathbf{x} \in \partial \Omega \tag{C.2}
\end{align*}
$$

where

$$
\begin{equation*}
\omega=\tau_{0} \omega_{0} \tag{C.3}
\end{equation*}
$$

Using separation of variable, we write $G(r, \theta)$ as

$$
\begin{equation*}
G=G_{0}(r)+\sum_{m=1}^{\infty}\left(G_{m}(r) e^{i m \theta}+c . c\right) \tag{C.4}
\end{equation*}
$$

where c.c refers to the complex conjugate of the term involving the summation. Substituting (C.4) into (C.1) and recalling the Newmann boundary condition, we obtain:

$$
\begin{gather*}
\frac{\partial^{2} G_{0}}{\partial r^{2}}+\frac{1}{r} \frac{\partial G_{0}}{\partial r}=1-\pi \delta\left(r-r_{0}\right) \delta(\theta), \quad G_{0} \text { bounded as } r \rightarrow 0, \quad G_{m}^{\prime}(1)=0  \tag{C.5a}\\
\frac{\partial^{2} G_{m}}{\partial r^{2}}+\frac{1}{r} G_{m}-\frac{m^{2}}{r^{2}} G_{m}+i m \omega G_{m}=-\pi \delta\left(r-r_{0}\right) \delta(\theta), \quad m>0, \quad G_{m} \text { bounded as } r \rightarrow 0, \quad G_{m}^{\prime}(1)=0 \tag{C.5b}
\end{gather*}
$$

For $m>0$, the homogeneous solution of (C.5b) may be written as

$$
G_{m}(r, \omega)=a_{m} I_{m}\left(c_{m} r\right)+b_{m} K_{m}\left(c_{m} r\right) ; \quad c_{m} \equiv \sqrt{-i \omega m}
$$

where $I_{m}(r)$ and $K_{m}(r)$ are m-th order modified Bessel functions of the first and second kind, respectively. Solving (C.5b) separately for $r<r_{0}$ and $r>r_{0}$, and applying appropriate continuity and jump conditions at $r=r_{0}$, we obtain the solution for $G_{m}$,

$$
G_{m}(r)=\left\{\begin{array}{cc}
\frac{1}{2}\left[-\frac{K_{m}^{\prime}\left(c_{m}\right)}{I_{m}^{m}\left(c_{m}\right)} I_{m}\left(c_{m} r_{0}\right)+K_{m}\left(c_{m} r_{0}\right)\right] I_{m}\left(c_{m} r\right), & 0<r<r_{0} \\
\frac{1}{2}\left[-\frac{K_{m}^{\prime}\left(c_{m}\right)}{I_{m}^{\prime}\left(c_{m}\right)} I_{m}\left(c_{m} r\right)+K_{m}\left(c_{m} r\right)\right] I_{m}\left(c_{m} r_{0}\right), & r_{0}<r<1
\end{array}, \quad c_{m} \equiv \sqrt{-i \omega m}, \quad m \neq 0\right.
$$

where $I_{m}^{\prime}\left(c_{m}\right)$ and $K_{0}^{\prime}\left(c_{m}\right)$ denote the derivatives of $I_{m}$ and $K_{m}$ evaluated at $c_{m}$, respectively. In a similar way, we find that the solution to (C.5) for $G_{0}(r)$,

$$
G_{0}(r)=\frac{r^{2}}{4}+\frac{2 r_{0}^{2}-3}{8}-\left\{\begin{array}{cc}
\frac{1}{2} \log r_{0}, & 0<r<r_{0} \\
\frac{1}{2} \log r, & r_{0}<r<1
\end{array}\right.
$$

Recall that

$$
R_{0}=G-S_{1}-S_{2}
$$

where

$$
\begin{equation*}
S_{1}:=-\frac{1}{2} \log \left|x-x_{0}\right| ; \quad S_{2}:=\frac{1}{4} \tau_{0} \frac{\partial x_{0}}{\partial t} \cdot\left(x-x_{0}\right) \log \left|x-x_{0}\right| \tag{C.6}
\end{equation*}
$$

To calculate $R_{0}$ and its gradient, we first expand the singular parts $S_{1}$ and $S_{2}$ in terms of their Fourier series, then take the limit $\theta \rightarrow 0, r \rightarrow r_{0}^{-}$. We have

$$
\begin{gathered}
S_{1}:=-\frac{1}{2} \log \left|x-x_{0}\right|=-\frac{1}{2} \log \left(r_{M}\right)+\frac{1}{2} \sum_{m \geq 1} \frac{\rho^{m}}{2 m}\left(e^{i m \theta}+e^{-i m \theta}\right), \text { where } r_{M}=\max \left(r, r_{0}\right), \quad \rho=\frac{\min \left(r, r_{0}\right)}{\max \left(r, r_{0}\right)} \\
\frac{1}{4} \tau_{0} \frac{\partial x_{0}}{\partial t} \cdot\left(x-x_{0}\right)=-\frac{\omega r r_{0}}{8}\left(i e^{i \theta}-i e^{-i \theta}\right) ; \\
S_{2}=\frac{\omega r r_{0}}{8 i}\left[\left(\log \left(r_{M}\right)+\frac{\rho^{2}}{4}\right) e^{i \theta}-\frac{1}{2} \sum_{m \geq 2}\left(\frac{1}{m-1} \rho^{-1}-\frac{1}{m+1} \rho\right) \rho^{m} e^{i m \theta}\right]+c . c .
\end{gathered}
$$

The function $F_{1}$ and $F_{2}$ defined through (6.81) are then expressed in terms of polar variables as

$$
F_{1}\left(r_{0}, \omega\right)=\left.\partial_{r} R_{0}\right|_{r=r_{0}^{-}}, \quad \theta=0 ; \quad F_{2}\left(r_{0}, \omega\right)=\left.\frac{1}{r_{0}} \partial_{\theta} R_{0}\right|_{r=r_{0}^{-}, \quad \theta=0}
$$

Differentiating with respect to $r$ and $\theta$ and then evaluating at $r=r_{0}^{-}$and $\theta=0$, we finally obtain the following expressions,

$$
\begin{align*}
F_{1}\left(r_{0}, \omega\right) & =\frac{r_{0}}{2}+\sum_{m \geq 1}\left(2 \operatorname{Re}\left(G_{m}^{\prime}\left(r_{0}^{-}\right)\right)-\frac{1}{2 r_{0}}\right)  \tag{C.7a}\\
r_{0} F_{2}\left(r_{0}, \omega\right) & =-\frac{\omega r_{0}^{2}}{4}\left[\left(\log \left(r_{0}\right)+\frac{1}{4}\right)\right]-2 \operatorname{Im}\left(G_{1}\left(r_{0}^{-}\right)\right)+\sum_{m \geq 2}\left(-2 m\left(\operatorname{Im} G_{m}\left(r_{0}^{-}\right)\right)+\frac{\omega r_{0}^{2}}{4} \frac{m}{m^{2}-1}\right) \tag{C.7b}
\end{align*}
$$

where

$$
\begin{align*}
G_{m}\left(r_{0}^{-}\right) & =\frac{1}{2}\left[-\frac{K_{m}^{\prime}\left(c_{m}\right)}{I_{m}^{\prime}\left(c_{m}\right)} I_{m}\left(c_{m} r_{0}\right)+K_{m}\left(c_{m} r_{0}\right)\right] I_{m}\left(c_{m} r_{0}\right)  \tag{C.7c}\\
G_{m}^{\prime}\left(r_{0}^{-}\right) & =\frac{c_{m}}{2}\left[-\frac{K_{m}^{\prime}\left(c_{m}\right)}{I_{m}^{\prime}\left(c_{m}\right)} I_{m}\left(c_{m} r_{0}\right)+K_{m}\left(c_{m} r_{0}\right)\right] I_{m}^{\prime}\left(c_{m} r_{0}\right), \quad c_{m} \equiv-i \sqrt{i \omega m} \tag{C.7d}
\end{align*}
$$

The Hopf bifurcation threshold derived in Proposition 4.1 corresponds to letting $r_{0} \rightarrow 0$. To establish the equivalence between the expression for $r_{0}$ in Proposition 6.1 and the threshold $\tau_{h}$, small in Proposition 4.1, we using the small-argument expansions for $K_{m}$ and $I_{m}$ to obtain the leading-order expressions,

$$
F_{1}(r, \omega)=\frac{r_{0}}{2}+2 \operatorname{Re}\left(G_{1}^{\prime}\left(r_{0}^{-}\right)\right)-\frac{1}{2 r_{0}} ; \quad r_{0} F_{2}\left(r_{0}, \omega\right)=-2 \operatorname{Im}\left(G_{1}\left(r_{0}\right)\right)-\frac{\omega r_{0}^{2}}{4}\left[\left(\log \left(r_{0}\right)+\frac{1}{4}\right)\right]
$$

Further simplifying, we obtain

$$
\begin{aligned}
& F_{1}\left(r_{0}, \omega\right) \sim \frac{r_{0}}{4}\left(-\operatorname{Im}\left(\frac{K_{1}^{\prime}(\sqrt{-\omega i})}{I_{1}^{\prime}(\sqrt{-\omega i})}\right) \omega+2-\frac{\pi \omega}{4}\right), \quad r_{0} \ll 1 \\
& F_{2}\left(r_{0}, \omega\right) \sim \frac{1}{4} r_{0} \omega\left(-\operatorname{Re}\left(\frac{K_{1}^{\prime}(\sqrt{-\omega i})}{I_{1}^{\prime}(\sqrt{-\omega i})}\right)+\frac{1}{2} \log (\omega / 4)+\gamma-\frac{1}{2}\right), \quad r_{0} \ll 1
\end{aligned}
$$

Setting $F_{1}=0$ and letting $r_{0} \rightarrow 0$, we obtain $\omega=\omega_{c}$ is the root of (4.51a). Then setting $F_{2}=0$ and recalling that $\omega=\omega_{0} \tau_{0}$, one recovers Proposition 6.1.
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